PACKMAN-Net: A Distributed, Open-Access and Scalable, Network of User-Friendly Space Weather Stations.

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THE PACKMAN (Particle Counter K-index Magnetic Anomaly) Instrument Space Weather on Earth



The observations acquired by PACKMAN will be used to provide **open access, real time information**, for:

- 1) education and public awareness of space weather phenomena;
- 2) to compare with Earth climate observations;
- 3) to provide real-time information of space weather variability for potential damage to infrastructures (telecommunications, power generation facilities, aviation, transport, etcetc.);
- 4) to monitor natural radiation sources at multiple environments;
- 5) to monitor the variability of the Pfotzer maximum height during different stages of solar activity and seasons and
- 6) This project may serve as a reference for future scalable networks where multiple instruments are deployed at different sites or conditions and with different initiation times, and where the informational value increases by adhering to a common PDS4 format and analysing the data in a concurrent way.



PACKMAN-E (Surface Earth observations)

### PACKMAN monitoring approach

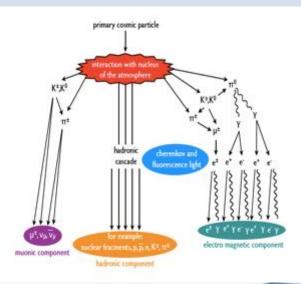
Solar Energetic Particles LEO (space weather: Near Earth Environment) PACKMAN-B (Balloon observations)

1. Solar Activity modulating: Galactic Cosmic Rays

2. Geomagnetic interaction: Incoming primary rays modulation and deflection

3. Atmospheric interaction: Secondaries generation, Cloud nucleation

4. Surface measurement: variation with total column of air, height, latitude and local magnetic field



Arctic balloon campaigns to monitor the variability of space radiation between 10 km and 40 km height. Complemented with surface measurements at multiple latitudes.

### Atmospheric and surface observations

### **Atmosphere**

- Orbiter based observations are available on-line, in almost real time (graphical).
- Ionosphere monitoring is available on-line (graphical)
- There is a lack of measurements of systematic radiation profiles in the lower layers of the atmosphere in the Arctic.

### Surface

- Only neutron counters (from galactic cosmic rays) are available on-line (graphical).
- K-index (geomagnetic variability) is monitored as a proxy for solar activity at multiple latitudes.
- It is more complex to have access to crude data in real time. These instruments are operated by scientific institutions.

<u>Purpose:</u> deploy network of PACKMAN with real time on-line data access, deploy equivalent instrument on balloon campaign with access on-line to the data. Compare time-variability sequence from satellite to surface, and weather/climate records.

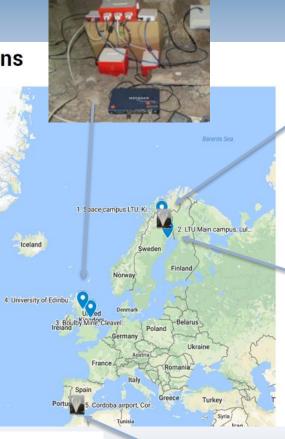
### **PACKMAN Locations**



United Kingdom

5. Cordoba airport, Cordoba, Spain

6. Esrange Space Center, Kiruna, Sweden







#### Initial deployment:

Through the initial deployment and testing phase, PACKMAN has demonstrated its operability at multiple latitudes and atmospheric heights:

- Space campus LTU, Kiruna, Sweden (67.84°N, 20.41°E, 390 m)
- LTU Main campus, Luleå, Sweden (65.62°N, 22.14°E, 15 m)
- Boulby Mine, Cleaveland, United Kingdom (54.56°N, 0.82°W, 93 m and -1.1 km)
- University of Edinburgh, United Kingdom (55.94°N, 3.19°W, 98 m)
- Cordoba airport, Córdoba, Spain (37.84°N, 4.84°W, 90 m to 27 km)
- Esrange Space Center, Kiruna, Sweden (67.88°N, 21.12°E, 328 m to 27 km)





Circumpolar long duration stratospheric balloon campaigns

1 week duration space weather stratospheric measurements.

Started from ESRANGE (Sweden) operated by SSC. HEMERA European call for instruments.







## User community

#### Global Positioning System (GPS)

Geomagnetic storms can impact the accuracy and availability of OPS by changing the ionorphere, the electrically marged layer of the atmosphere a GPS signal must pass through from synellite to provind receiver. The ionorphere is the largest source of error in GP positioning and navigation. These ionospheric disturbances are lower our sent but can become severe during geomagnetic storms resulting in

range errors in excess of 100 feet, or ever resulting in loss of lock on the GP1 signal entirely. These errors can have signafican impacts on precision uses of GP5 such at navigation, agriculture, oil drilling surve) ing, and timing.



#### Satellite Operations

There are thousands of satellites in orbit around Earth with television and radio, communications, meteorology, national defense, and much more. Space weather can affect these satellites in many ways. Solar radiation storms can cause spacecraft orientation problems by interfering with star trackers and by causing errors or damage in electronic devices. Geomagnetic storms can create a hazardous charging environment for satellites resulting in damaging electrostatic discharge, much like touching a door knob and getting that spark on a dry winter day. Geomagnetic storms also cause heating of the atmosphere, essentially causing it to expand, which results in more drag or slowing down of an orbiting satellite. In a worst case, space weather can cause the

#### Space Operations

satellite to fail.

Astronauts and their equipment in space are bombarded with charged particle radiation. This radiation causes tissue or cell damage in humans. Space weather and solar radiation storms are of particular concern for activities outside the protection of Earth's atm ophere and magnetic field.

## Space Weather mpacts on Earth

Electrons accelerated in the tail of the magnetosphere travel down the magnetic field lines.

Electrons collide with the apper atmosphere 50 to 300 miles above Earth.

Electrons exchange enorgy with the atmosphere of end the atmosphere atoms of molecules to higher energy level. When the atoms and molecules folax back to energy levels, the cleans the orienty in the form of light.

#### Aurora

The Aurora Borealis (Northern Lights) and Aurora Australis (Southern Lights) are the result of electrons colliding with Earth's upper atmosphere. The electrons are energized through acceleration processes in the downwind tail (nightside) of the magnetosphere. The accelerated electrons follow the magnetic field of Earth down to the polar regions where they collide with oxygen and nitrogen atoms and molecules in Earth's upper atmosphere. In these collisions, the electrons transfer their energy to the atmosphere, thus exciting the atoms and molecules to higher energy states. When they relax back to lower energy states, they release their energy in the form of light. The aurora typically forms 50 to 300 miles above the ground. Earth's magnetic field guides the electrons such that the aurora forms two ovals approximately centered at each magnetic pole.

#### THE COLORS OF THE AURORA

 Deep red from high altitude atomic oxygen

Magenta from high altitude molecular nitrogen in sunlight

Greenish yellow from lower altitude atomic oxygen

Magenta from low altitude molecular nitrogen (not shown in the picture)

#### Aviation

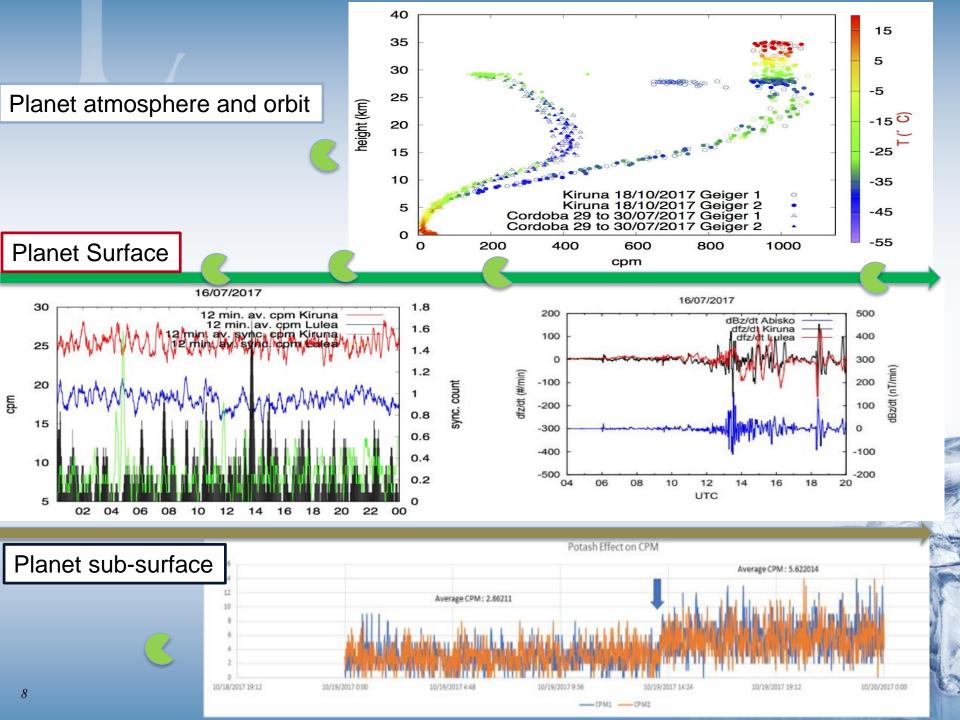
Aircraft use fligh frequency (HF) radio communication to stay in touch with ground controllers in temote areas such as over the occars or over the poles. Solar flares can thack out the use of HE on the dayside of carth and solar radiation storms can 'black out' use of HE near the poles, impacting the aircraft's ability to stay in touch with the ground, limpacts to GPS systems can also significantly affect airline operations.

#### Power Grids

Geomanetic storms result in electric currents in the manetosphere and ionosphere as the area shaped by Earth's magnetic field is compressed and disturbed. The disturbed conditions create additional numerits in long conductors on the ground such as overhead transmission lines or long pipelines. In the most extreme cases, these currents can cause voltage instability or damage to lower estem components potentially roluting in temporary service disruptions, or zero a videspread power outage.

NOAA Education www.education.noaa.gov NOAA Space Weather Prediction Center www.spaceweather.gov

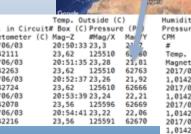
Image source: Aurora Boreally, altern from the International Space Station in April of 2012.



### Open access to planetary data

	AERONAUTICS E ADMINISTRATION		+ NASA Homepage + NASA en Español + Contact NASA	•		
2.10		ciences Node				
HOME DATA	AND SERVICES TOOL	S ABOUT US CI	ONTACT US SITE			
Services	Mars Science Lab	oratory (MSL)	What's New			
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Mars		y (MSL) rover, Curiosity, landed on	sols 360-449 have bee			
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Moon Earth	DAN - Dynamic Albedo of M	Neutrons DAN Archive	derived data acquired-			
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Help for Data Users Help for Data Reviewers	Mastcam - Mast Camera		August 30, 2013. 5/5/	-		
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Because of our heritage from the and the **REMS** Mars Science Laboratory (NASA) and the HABIT ExoMars mission (ESA): the PACKMAN data will have a similar structure as the PDS4 protocol that is used by NASA/ESA for planeta ation.



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## PACKMAN data volume

- Data storage: 132 Kilobytes/Day
- PACKMAN data have 11 columns of measurement that start with a time stamp and latitude/coordinate/elevation information.
- Every new line corresponds to a new time stamp, with measuring interval of 1 minute in UTC.
- All PACKMAN nodes provide observations with the same time stamp, and same observation cadency.
- 365 days of continuous measurements/year.
- Scalable network:

10

- Internal nodes uploading daily.
- External nodes uploading at blocks of 3 months (pre-acceptance required).

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## PACKMAN data requirements

# Data uploading (Instrument)

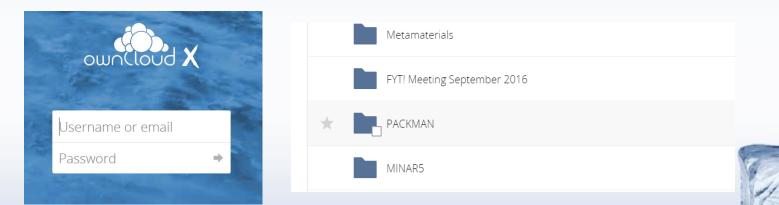
- Open access –"preauthorized" users- data storage.
- Scalable (over time more nodes).
- Automatic PDS4
   labelling

Data downloading (Science/ Infrastructures/ Education/ Outreach)

- Global: open access
- Online tools for quick preliminary analysis (online representation, online fast processing) and comparison of multiple nodes-data.
- Accessibility to multiple
   simultaneous instrument
   observations.

# Open-access pre-authorized data uploading

 Owncloud: Private Owncloud server for third party users to upload their PACKMAN Data automatically using the official Owncloud client.



'HNOLOG'

 Automation tools: Using Python scripts to create the repositories and put the data uploaded to owncloud and check new files everyday.

# Labelling PDS4

 Label creation: Use of MakeLabelsPDS4 tool created by <u>Dan</u> <u>Scholes (PDS GeoSciences Node)</u> using Excel to populate the PDS4 XML fields over XML Templates

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1	spectrum_id	file_name	create_date	lab_name	lab_desc	instr1_name	instr1_id	instr2_name	instr2_id	meas1_date	
			Creation Date	Laboratory			Instrument #1	Instrument #2	Instrument #2	Measurement #1 Date	
2	Spectra ID	File Name	(YYYY-MM-DD)	Name	Laboratory Description	Instrument #1 Name	ID	Name	ID	(YYYY-MM-DD)	
3	spectra_id1	spectra1.dat	2000-01-02	Lab Name 1	University 1 laboratory	Lab 1 Bidirectonal Reflectance Spectrometer	bdrs.lab1			2000-01-02	
4	spectra_id2	spectra2.dat	2001-01-02	Lab Name 2	University 2 laboratory	Lab 2 Bidirectonal Reflectance Spectrometer	bdrs.lab2			2001-01-02	
5	spectra_id2	spectra3.dat	2003-01-03	Lab Name 3	University 3 laboratory	Lab 3 Bidirectonal Reflectance Spectrometer	bdrs.lab3			2003-01-03	
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### **PACKMAN Data File**

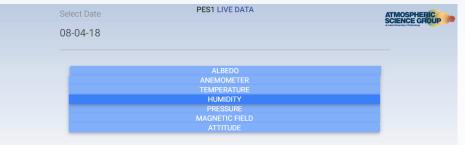
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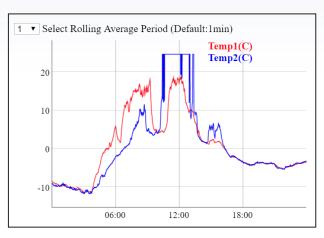
# Online quick analysis: cross-instrument comparison (multi latitud)

- Online data visualization
- (https://atmospheres.research.ltu.se/pes1)



🔟 PLOT

Online plotting of different variables using → **DYGRAPHS** API (www.dygraphs.org)





# Summary: PACKMAN-Net

- This is an example of planetary instrumentation relevant data (Space Weather) in our planet which needs a dedicated scalable data archiving and processing architecture that adheres to the PDS4 standard.
- Furthermore, this project may be used to benchmark the design of archiving, scalable, networks of future planetary instrumentation observations of the Moon or mars or Earth orbiters.
- This project will bridge the gap between society and research, adding **new stake holders** such as teaching institutions (high schools, universities), or industry and infrastructures representatives.